John Archibald Wheeler, 1952-1976: Black Holes and Geometrodynamics



Kip S. Thorne, Caltech [relying heavily on Charles W. Misner, U. Md.] American Physical Society, Denver, 2 May 2009

Relativity in 1952

- A few exact solutions: Schwarzschild, Reissner-Nordstrom, Friedman, ...
- Linearized approximation: gravitational waves skepticism over their physical reality
- Post-Newtonian equations of motion
- Contact with observation:
 - Expansion of universe
 - one precision test: perihelion shift of Mercury
 - two crude tests: light deflection, gravitational redshift
- Moribund; backwater of physics; little activity since 30s



- 1953: Wheeler, emerging from H-bomb design work, embarked on quest to understand general relativity; curved spacetime
- First step:
 "If you would learn, teach"
- Wheeler led a "revolution";
 → major field of fundamental physics & tool for astrophysics



955 - GEON: Mass without Mass

- Gravitational Electromagnetic "particle"
- Leakage; collective instability so can't exist in Nature

BUT:



- Hinted at richness that nonlinearities can produce in curved spacetime
- In hands of Wheeler's students \rightarrow powerful tools:
 - Brill & Hartle: 2-lengthscale expansions
 - Isaacson: Stress-energy tensor for Grav'l Waves

957 - Wormhole: Charge without Charge

- Building on: Flamm 1916;
 Weyl 1924; Einstein&Rosen 1936
- "Radical Conservative-ism"
 - adhere to well-established physical law
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- Pinch-off
- Quantum foam
- Wheeler DeWitt Eqn



Gravitational Waves

- I956: Joseph Weber: Wheeler's "postdoc"
 - Weber & Wheeler, RMP (1957) - on the physical reality of gravitational waves





Geometrodynamics

- The dynamics of curved spacetime
- Give initial 3-geometry of space, subject to constraints.
- Evolve the geometry forward in time
- Mathematical formulation: Arnowitt, Deser & Misner -1959, 1960; York 1971 ff
- Wheeler's vision: numerical relativity
 - solutions to constraint equations for multiple wormholes: Misner 1960; Brill & Lindquist 1963





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 - cut off at throats get black holes
 - 111 citations since 2000



Geometrodynamics Today





What else? What between neutron stars & white dwarfs?



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How to Discover a Neutron Star?

Wheeler: 1966, Annual Reviews of Astronomy & Astrophysics

- Crab Nebula: remnant of a supernova explosion
- Should be a Neutron Star inside
- How is the Crab energized?
- Wheeler's speculation: a spinning, magnetized neutron star!









The Fate of Massive Stars

Wheeler's Conclusion, 1957



- Unacceptable answer! CRISIS for physics!!
- Speculation: Perhaps a "fiery marriage" of general relativity and quantum theory (Quantum Gravity) will take over, and replace the singularity by something else
- The Issue of the Final State





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 Principle of

Radical

Conservative-sm

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 Wheeler resisted. Confrontation at June 10, 1958 - Solvay Conference, Brussels





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*Perhaps most of all: Beckedorff & Misner elucidate

Oppenheimer & Snyder



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- **Oppenheimer & Snyder**
- ★1963: Quasars Discovered
- **★** First Symposium on Relativistic Astrophysics
- ★ Wheeler adopted the name **Black Hole** (1968)
- \star Continued to focus on the mystery of the singularity
- ★Until 1964, speculated information leaks out



Dopenheimer:

1964: Information Leaking Out of a Black Hole John and Kip - Writing *Gravitation Theory and Gravitational Collapse*







High Island, Maine

1964: Information Leaking out of Black Hole



Hawking Radiation before Hawking

1964: Information Leaking out of Black Hole



The Stability of Black Holes

1956: Tullio Regge and John Wheeler

PHYSICAL REVIEW

VOLUME 108, NUMBER 4

NOVEMBER 15, 1957

Stability of a Schwarzschild Singularity

TULLIO REGGE, Istituto di Fisica della Università di Torino, Torino, Italy

AND

JOHN A. WHEELER, Palmer Physical Laboratory, Princeton University, Princeton, New Jersey (Received July 15, 1957)

It is shown that a Schwarzschild singularity, spherically symmetrical and endowed with mass, will undergo small vibrations about the spherical form and will therefore remain stable if subjected to a small nonspherical perturbation.

I. INTRODUCTION AND SUMMARY

S CHWARZSCHILD found long ago the solution of the Einstein equations for the metric around a fixed spherically symmetrical center-of-mass:

$$ds^{2} = -(1 - 2m^{*}/r)dT^{2} + (1 - 2m^{*}/r)^{-1}dr^{2} + r^{2}(d\theta^{2} + \sin^{2}\theta d\varphi^{2}) = g_{\mu\nu}dx^{\mu}dx^{\nu}, \quad (1)$$

with $x^0 = T$, $x^1 = r$, $x^2 = \theta$, $x^3 = \varphi$. Here the quantity

$$m^{*}(cm) = Gm/c^{2} = (0.74 \times 10^{-28} cm/g)m(g)$$
 (2)

field in the usual way of writing these equations. However, the equations can be rearranged³ in such a way as to bring into evidence an additional production of gravitational field by the stress energy tensor of the gravitational field. On this account the geometrized mass, m^* , is not uniquely determined by the geometrized charge, q^* ; it only follows that m^* is no less than q^* . (5) One can therefore think of the field energy—or the mass and stress that goes with it—as in equilibrium under its own gravitational attraction.

We have equilibrium but is it stable? A sphere of

- My first meeting with Wheeler: June, 1962
- Wheeler's mentoring style



• Lectures: Tours de Force



Bill Unruh's recollections





Bob Geroch's recollections





David Sharp recollections





Misner, Thorne & Wheeler

GRAVITATION

Charles W. MISNER Kip S. THORNE John Archibald WHEELER

1973



